



Antarctic Survey Telescopes (AST3) Remote Sky Survey Strategy

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On behalf of:
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Overview



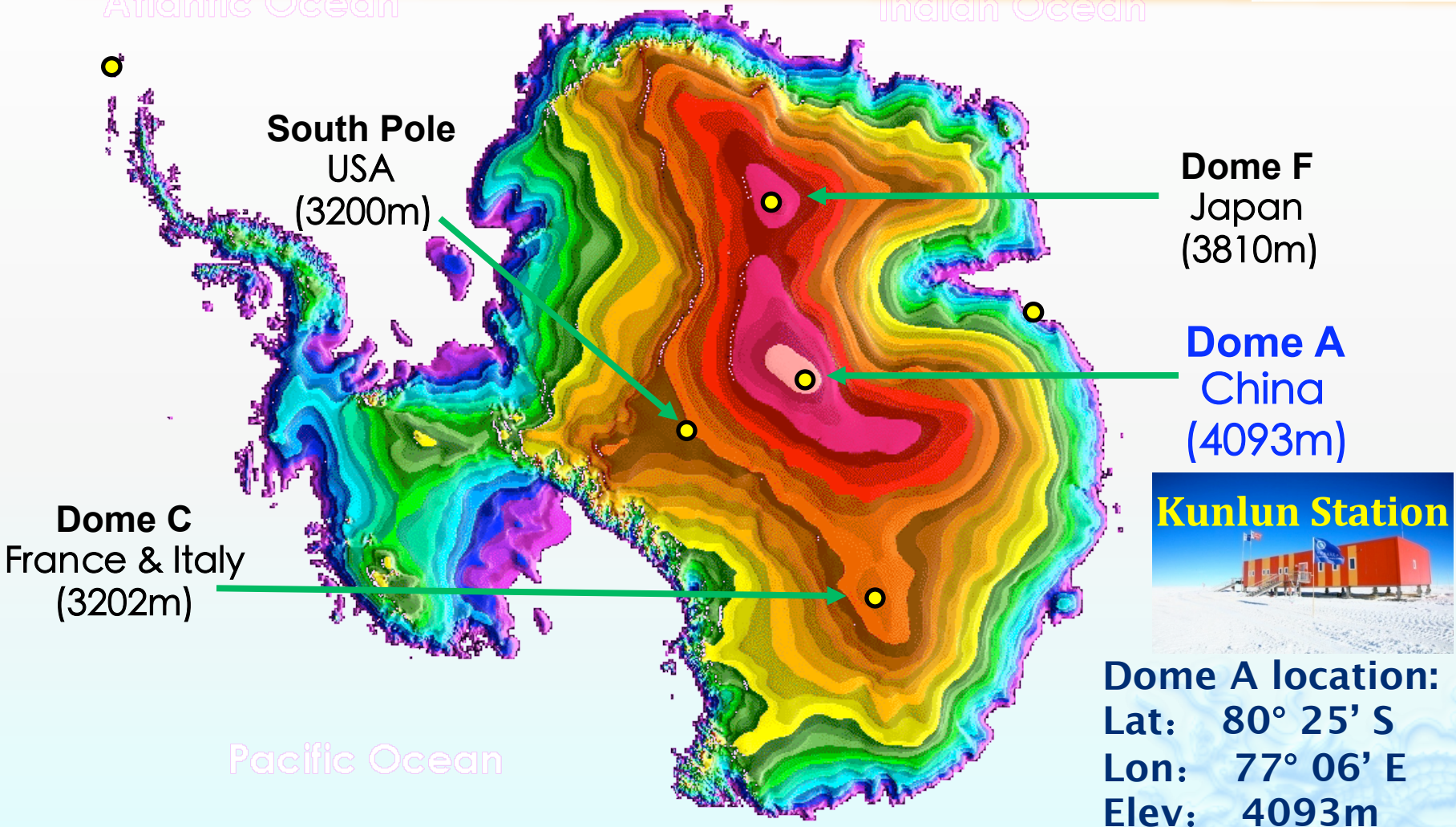
- Dome A & AST3
- Sky Survey Strategy
 - Sky Background
 - Extinction
 - Aurora & Cloud
 - Cadence
 - Movement of telescope
- Conclusions

Dome A (南极冰穹A)



Atlantic Ocean

Indian Ocean



South Pole
USA
(3200m)

Dome F
Japan
(3810m)

Dome A
China
(4093m)

Dome C
France & Italy
(3202m)

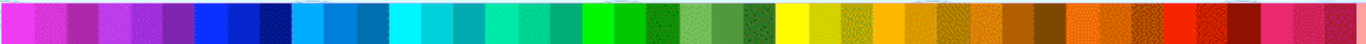


Dome A location:
Lat: 80° 25' S
Lon: 77° 06' E
Elev: 4093m

0

Elevation in meters

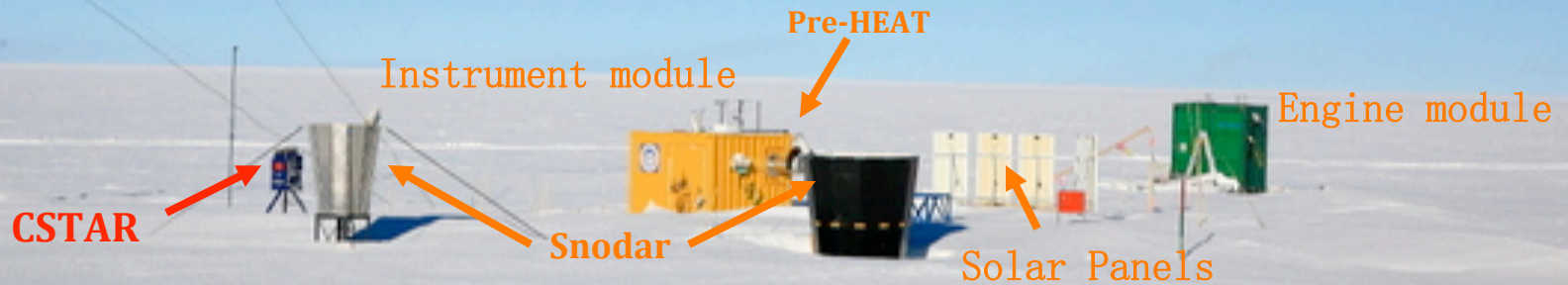
4000



Advantage of Dome A



- **Continuous observation during the polar night**
- **Low temperature (-70°C): infrared observation**
- **Wind speed (2.5m/s), stable atmosphere, low boundary layer(Snodar, AWS)
==> Good seeing**



- **Transparency, dark background (CSTAR)**
- **PWV<75 μ m: THz observation(Pre-HEAT, FTS)**
- **Cheaper than Space Mission**

Disadvantage of Dome A

- Difficult to access : Data transmission, maintenance, ...
- High requirement for the equipments : temperature, power consumption, ...

Not a problem until now and foreseeable future.



Antarctic Survey Telescopes (AST3)



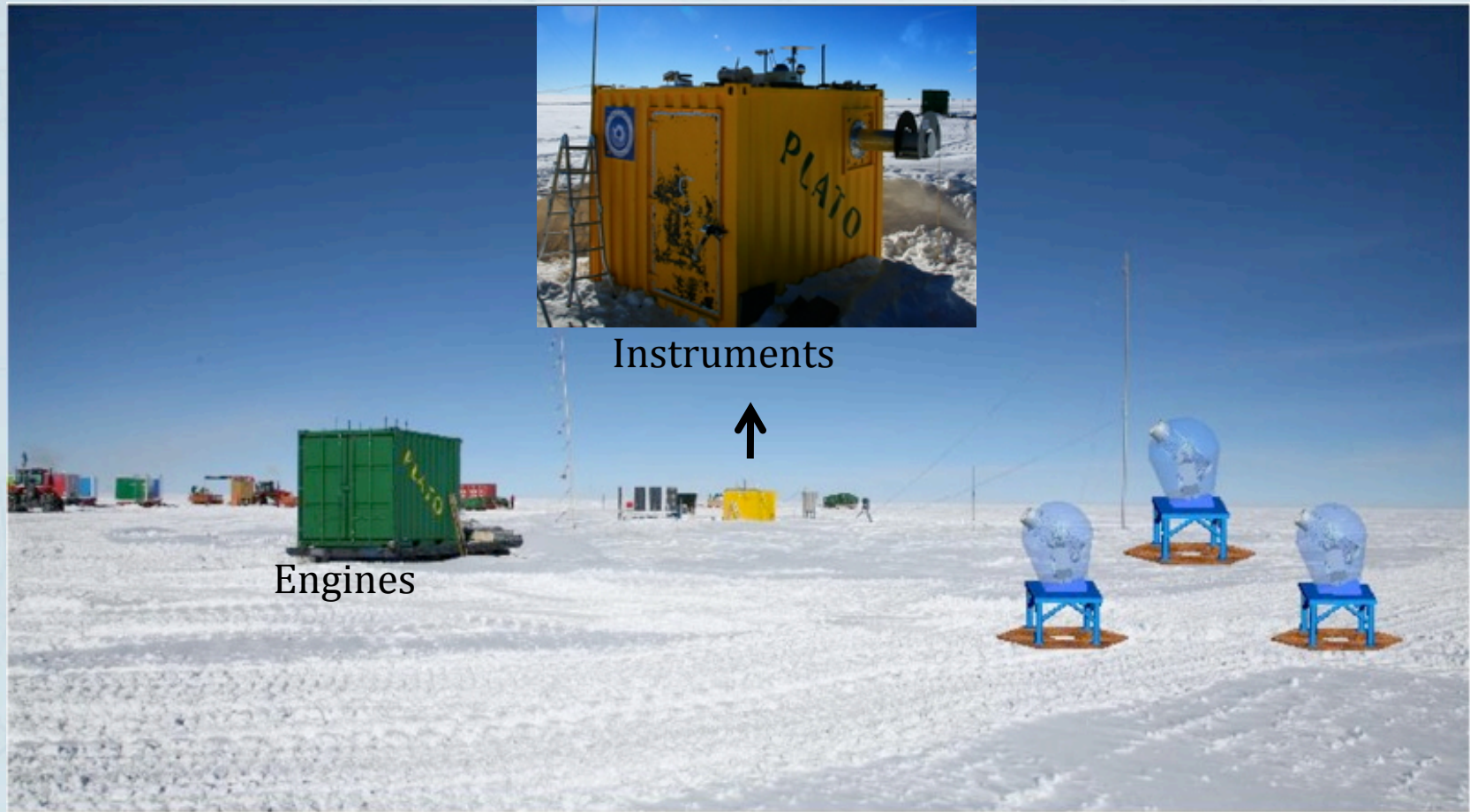
- Clear aperture: 500mm, Primary mirror: 680mm
- Wave band: 400nm-900nm (g, r, i filter for 3 telescopes)
- Field of view: 4.3 square deg
- Scale: 1 arcsec/pixel
- Image quality: 80% light energy encircled in 1 pixel
- Total optical length: 2.4m
- CCD: STA1600 , 9micron /pixel, 10K×10K: 5KX10K



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<http://aag.bao.ac.cn>

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Antarctic Survey Telescopes (AST3)



Instruments



Engines

Antarctic Survey Telescopes (AST3)



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Sky Survey Fields (Supernovae)

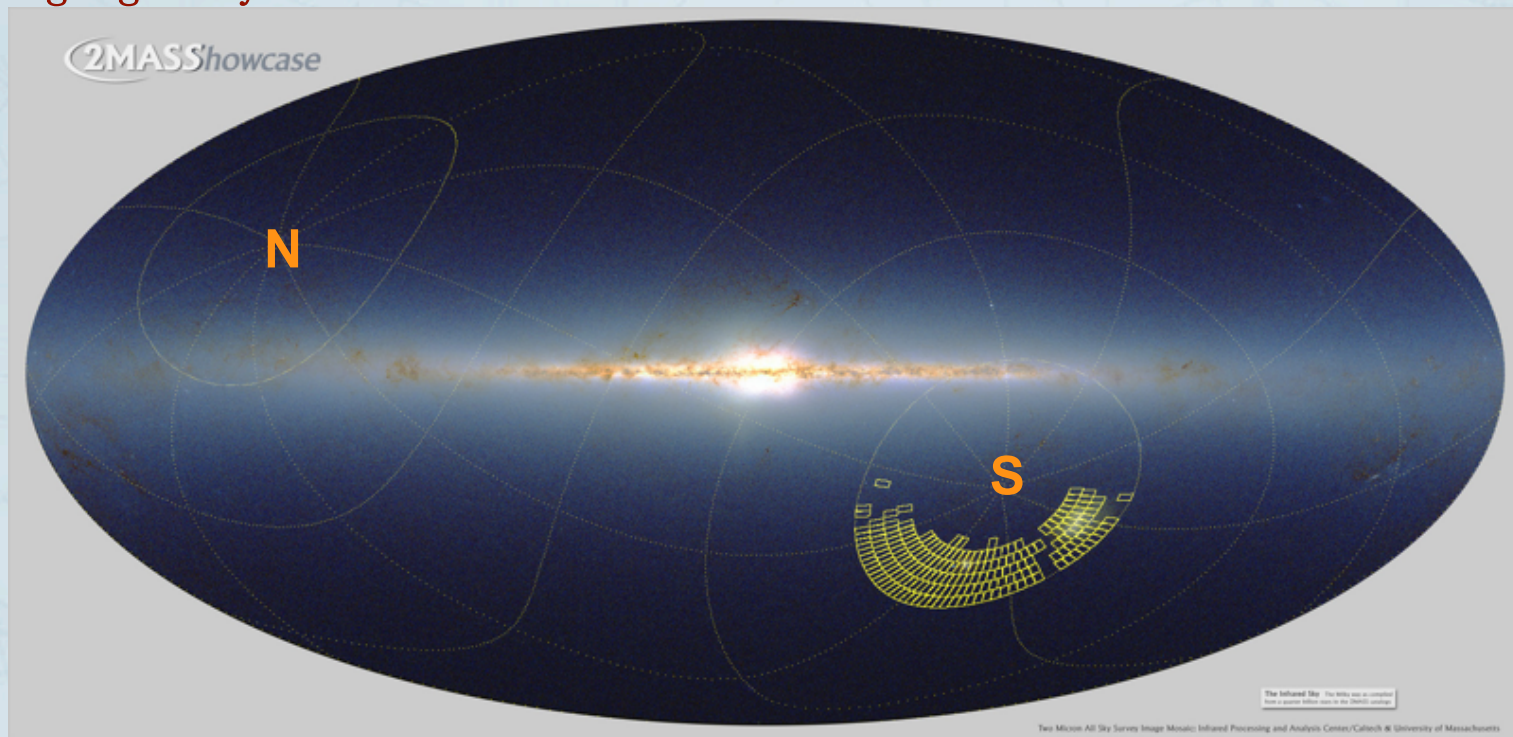


Criterion:

1. High galactic latitude
2. Low extinction
3. Large galaxy number



- 2000 sq degree
- ~500 fields



Credit to: Juncheng CHEN 、 Xiaofeng WANG (Tsinghua)

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Sky Survey Strategy(SSS)



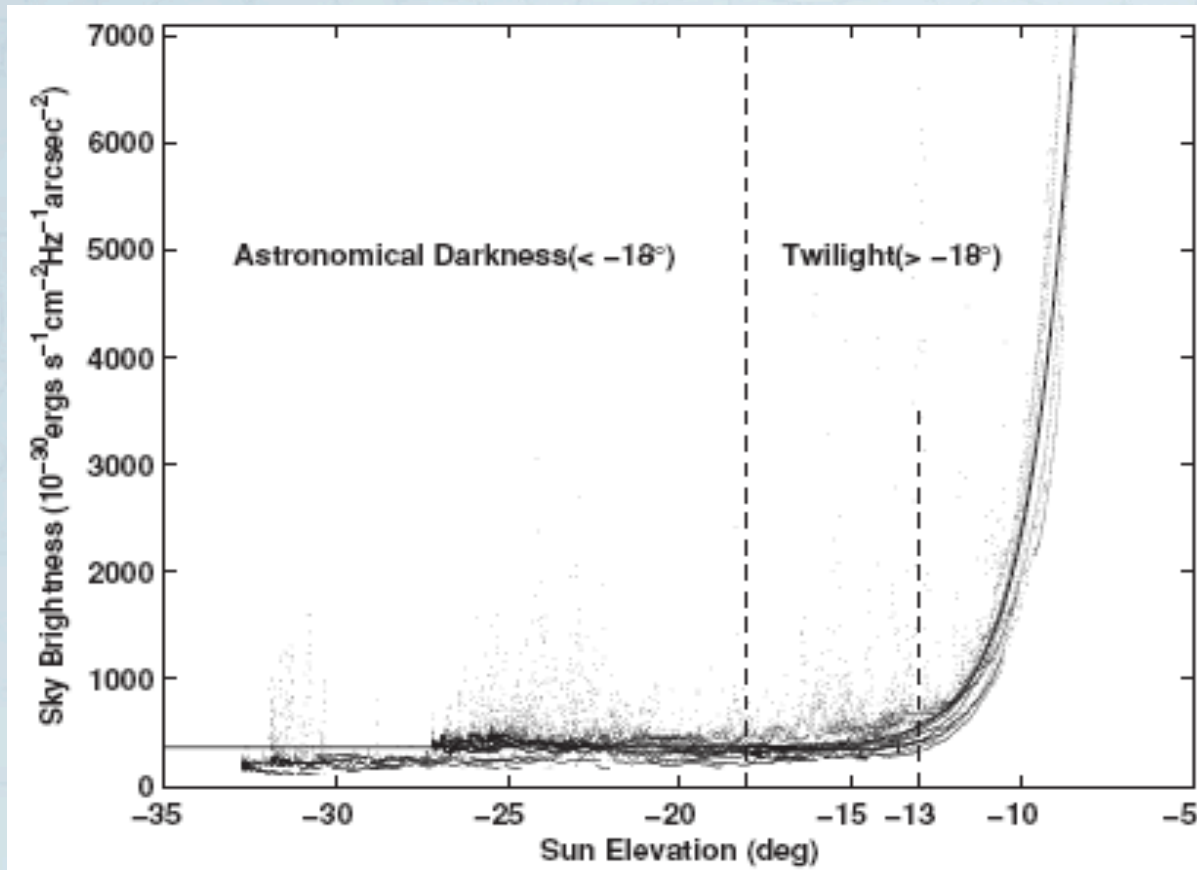
Select the best field to observe:

- **Dark sky background;**
 - Sun altitude (twilight)
 - Moon altitude & azimuth
 - Lunar phase
- **Extinction;**
- **Aurora、 cloud (possibility);**
- **Cadence;**
- **Movement of the telescope.**

SSS–Twilight



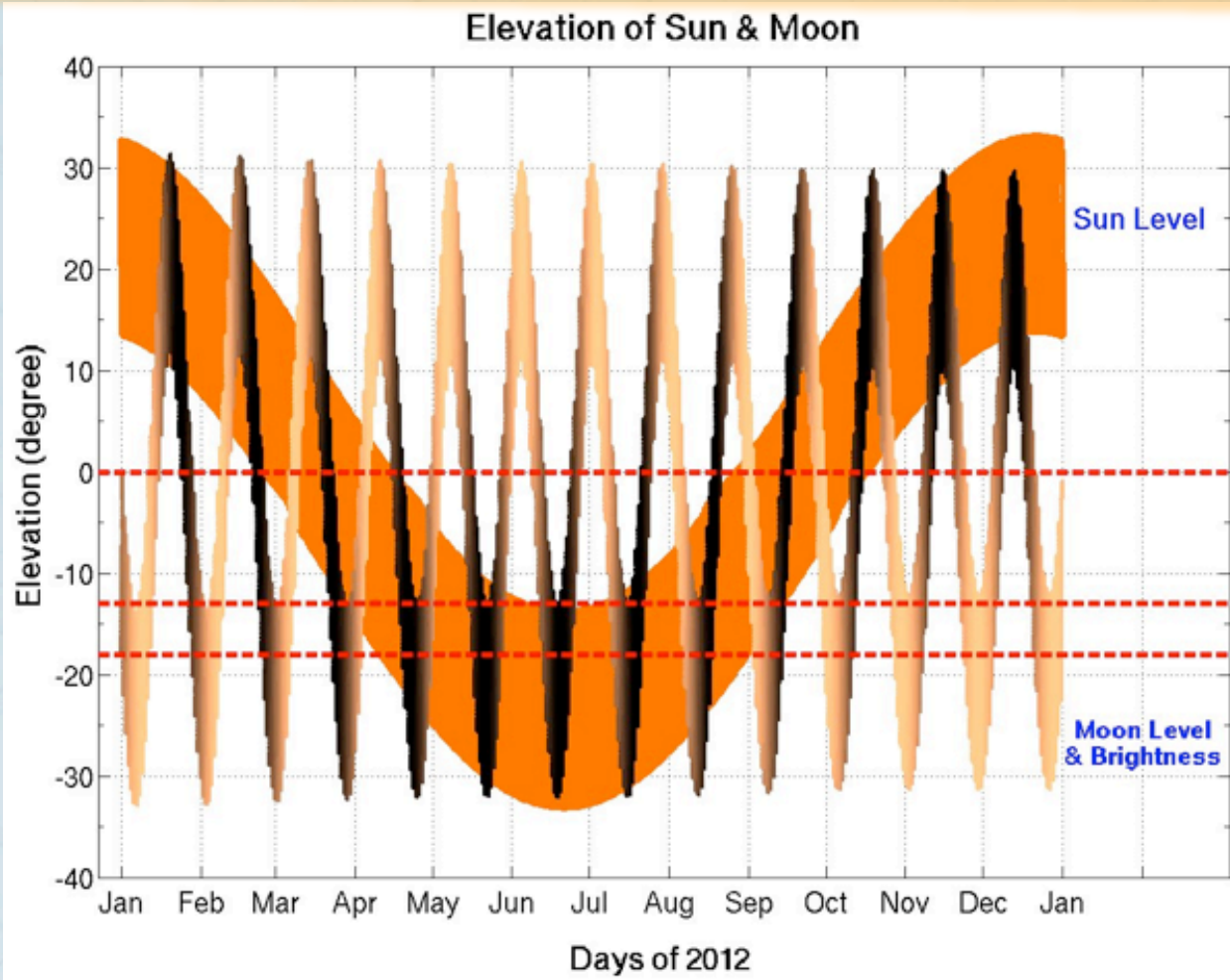
CSTAR result



The Sun altitude of -13° below the horizon is suitable twilight at Dome A

Zou et al. AJ, 2010

SSS - Altitude of the Sun



地平线

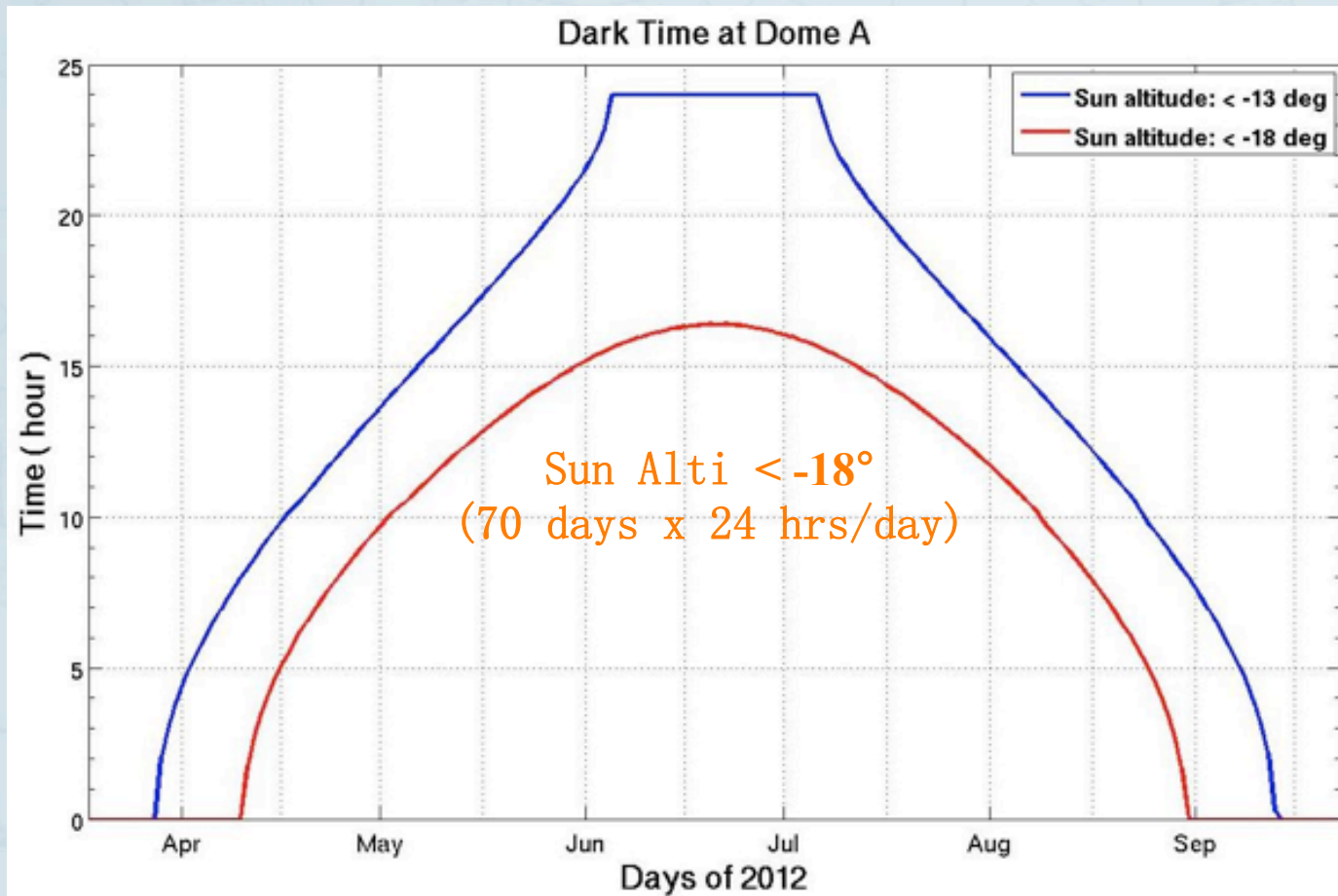
-13°

-18° twilight

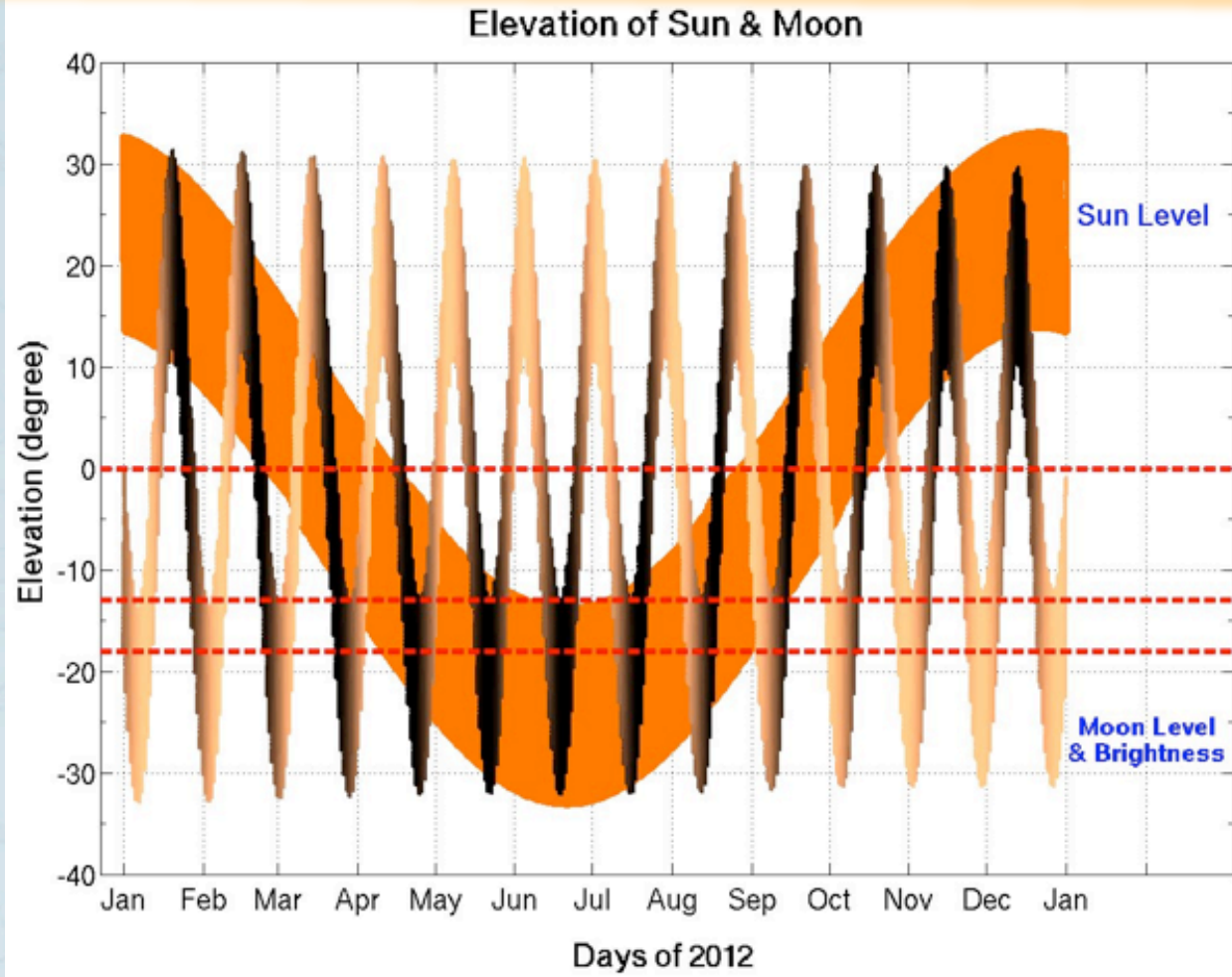
SSS– Total Observation Time



Sun Alti $< -13^\circ$ (109 days x 24 hrs/day)



SSS - Altitude of the Moon



地平线

-13°

-18° twilight

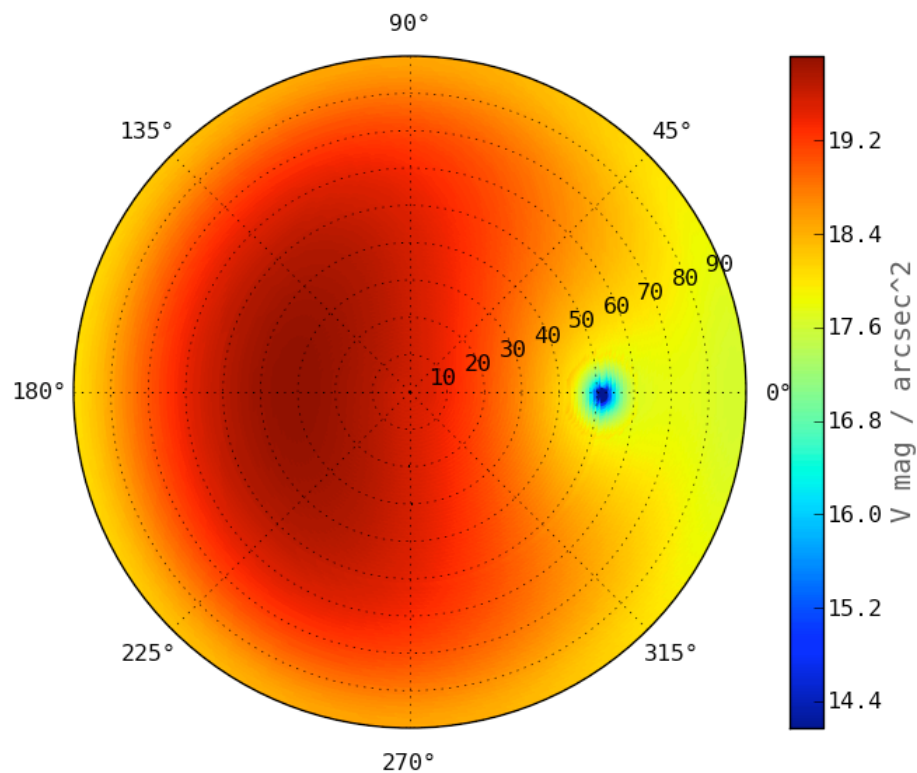
The largest altitude of the Moon in 2012: $< 35^\circ$

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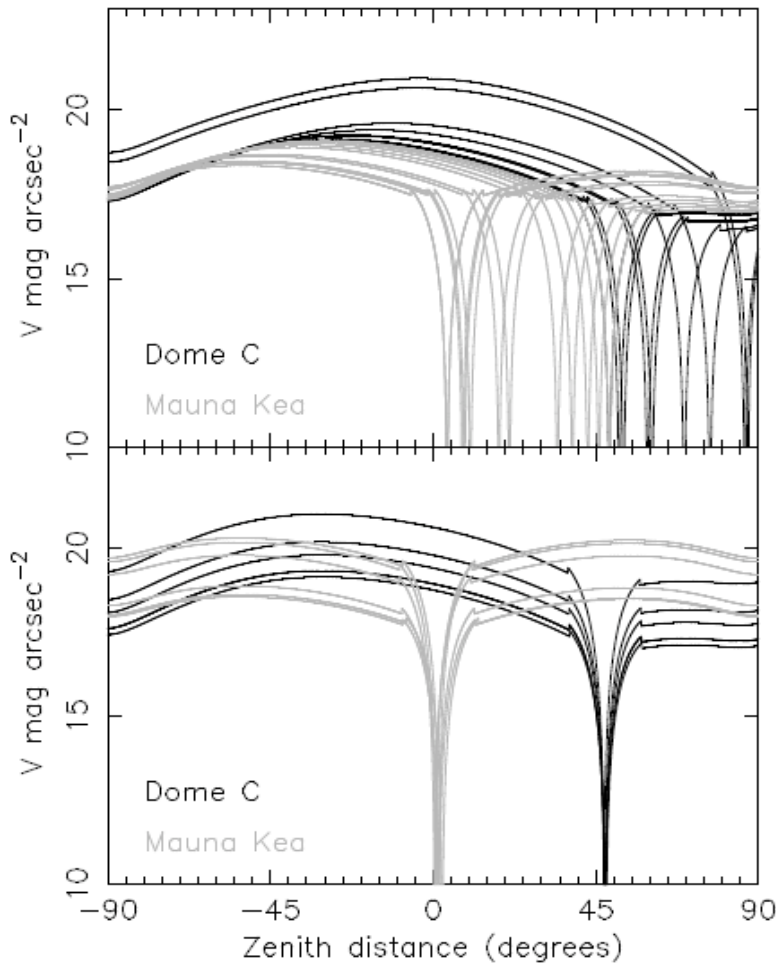
SSS– Moon Azimuth



Zenith distance @ Dome A = 50°
(the worst @ Dome A)

Model :
Krisciunas & Schaefer, 1991
Credit to: Geoffrey Sims

Moon Contribution to background



Kenyon & Storey, 2006

Full Moon reaches its highest altitude in each month (in 2005). Each curve represents a different month.

Moon reaches its highest altitude in each month (in 2005), regardless of phase.

Model : Krisciunas & Schaefer, 1991

Sky Survey Strategy



Select the best field to observe:

- Dark sky background;
- Low extinction (air mass) ;
 - Zenith distance $< 30^\circ$
- Aurora、cloud (possibility);
- Cadence;
- Minimize the movement of the telescope.

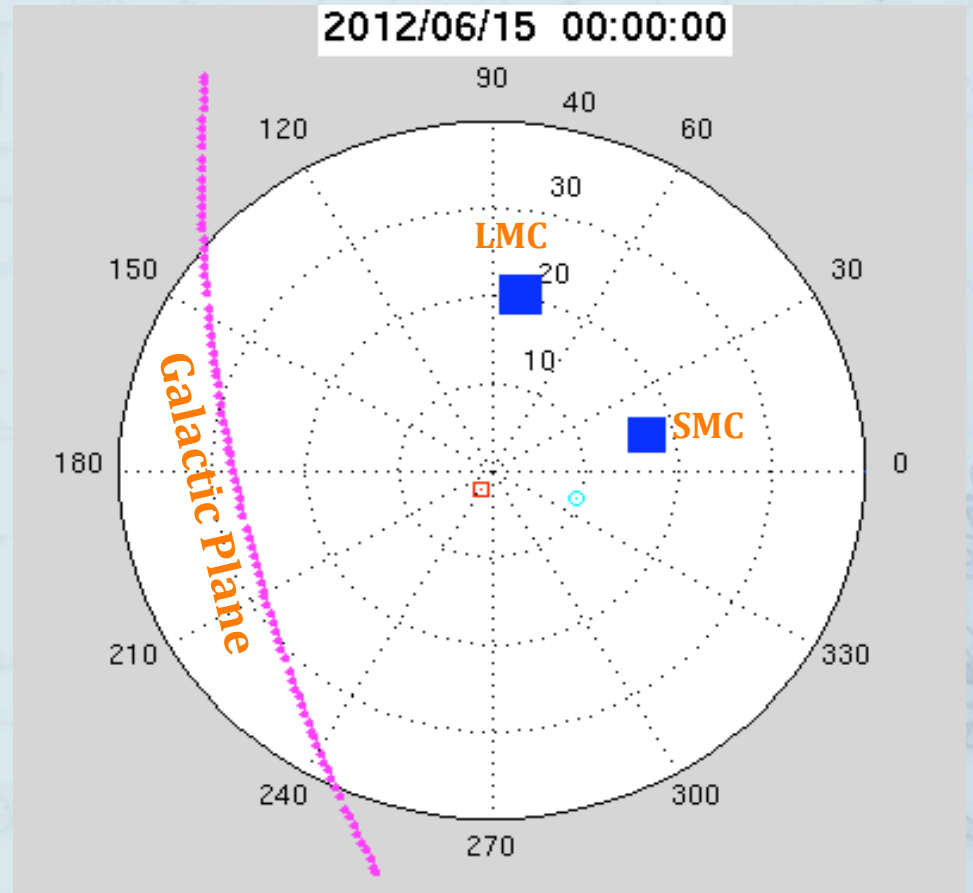


HRCAM

SSS - Simulation




- Galactic latitude $> -25^\circ$
- Sun Alt $< -13^\circ$
- Zenith distance $< 30^\circ$
- Minimize the movement



Conclusions



➤ Good Sky Survey Strategy  More Data (> 10T)

➤ CSTAR & AST3  Center for Information
&
Computing, NAOC, CAS

Thank CVO & LOC !